BUCKNELL UNIVERSITY Astronomy 101

Third Hour Exam

2005 November 16

This exam will be scored on a 100 point scale and has three parts:

- 1) a multiple choice section containing 6 questions each worth 4 points,
- 2) a short answer section containing 3 questions each worth 12 points, and 3) a problem section containing 2 problems worth 20 points each.

Below are some relationships you might find useful:

for a circle:

for a sphere: for an ellipse:

Kepler's Third Law:

acceleration:

for circular orbits:

gravitational acceleration:

for constant acceleration:

circumference = 2π radius

 $area = \pi radius^2$

volume = $4/3 \pi \text{ radius}^3$

eccentricity = f/a

 $(Period)^2 = (semi-major axis)^3$

acceleration = change in speed per time

or $a = \frac{\Delta v}{\Delta t}$

speed² = radius x acceleration

acceleration = $\frac{G \times mass}{radius^2}$ distance = 1/2 acceleration x time²

or $d = 1/2 \text{ a } t^2$

gravitational acceleration at the Earth's surface: 9.8 m/s²

The lens equation:

Planck's constant:

the speed of light:

the speed of sound:

radius of the Earth:

 $\frac{1}{focal\ length} = \frac{1}{source-lens\ distance} + \frac{1}{lens-image\ distance}$

wave speed = frequency x wavelength (i.e., $v = \nu \times \lambda$)

photon energy:

Wien's Law:

 $E_{photon} = \frac{h c}{\lambda}$

 $\lambda_{peak} = \frac{3.0 \times 10^6 \text{K nm}}{T}$

 $\lambda = 400 - 700 \text{ nm} (1 \text{ nm} = 10^{-9} \text{ m})$

 $G = 6.67 \times 10^{-11} \text{ m}^3 / \text{kg s}^2$

 $h = 6.63 \times 10^{-34}$ Joule seconds

 $6.379 \times 10^6 \text{ m}$

 $1.496 \times 10^{11} \text{ m}$

 $c = 3.00 \times 10^8 \text{ m/s}$

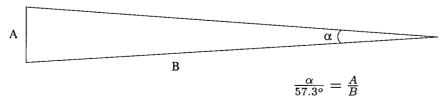
330 m/s

The Observer's Triangle Relation:

the visible wavelength range:

1 Astronomical Unit (A.U.)

Newton's gravitational constant:



Multiple Choice Questions (4 points each)

MC1. Double-rimmed craters are found on some planetary surfaces, caused by

- (a) an impactor that encountered a surface with material of different densities.
- b) multiple impacts at the same location.
- c) pyroclastic flow of molten rock during the impact.
- d) erosion of an originally single-rimmed crater.

MC2. The moon Titan has a thick atmosphere, while the planet Mercury has virtually none, because

- a) Titan is more massive than Mercury.
- b) the Sun's gravity stripped Mercury of its atmosphere.
- c) the cold gas in Titan's atmosphere is moving more slowly than the hot gas in Mercury's atmosphere.
- d) there's more gas in the outer solar system, so Titan was more easily able to acquire an atmosphere.

MC3. The Magellan spacecraft was able to "see" through the clouds in Venus' atmosphere and map the planet's surface by imaging

- a) ultraviolet light.
- b) visible light.
- c) infrared light.
- d)radio light.

d) Mars

e) Titania

e) sound waves.

MC4. Both Earth and Io have active volcanoes. At what other location in the solar system has active vulcanism been observed?

a) Europa	Must be current for
b) Titan c) Triton	us to see it

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MC5. Which of the following statements regarding Saturn's rings is not true?

- a) They are composed of a variety of ices.
- (b) They are likely the material left over from the formation of Saturn and its moons.
- c) They extend from the outermost reaches of Saturn out to Saturn's Roche limit.
- d) They are only a few hundreds of meters thick.
- e) They orbit under the influence of Saturn's gravity.

MC6. We infer that the surface of Venus is geologically young (i.e., 500 million years or so old) based on the fact that

- a) volcanoes are currently active on the surface.
- b) fault lines indicate plate tectonic motion.
- c) the atmosphere has a density 100 times higher than that on Earth.
- d) the lithosphere is very thick.
- (e) there are a small number of craters on the surface.

Short Answer Questions (12 points each)

SA1. Should Pluto be considered a planet? Explain why, and provide at least two pieces of hard evidence supporting your claim. (Note: I don't care which side of this argument you take — I'll grade your response based on how well you support your argument.)

- orbits the Sun

- has a moon

- is differentiated

- was first kBO discovered

and remained the only

such object for nso years

- has an atmosphere

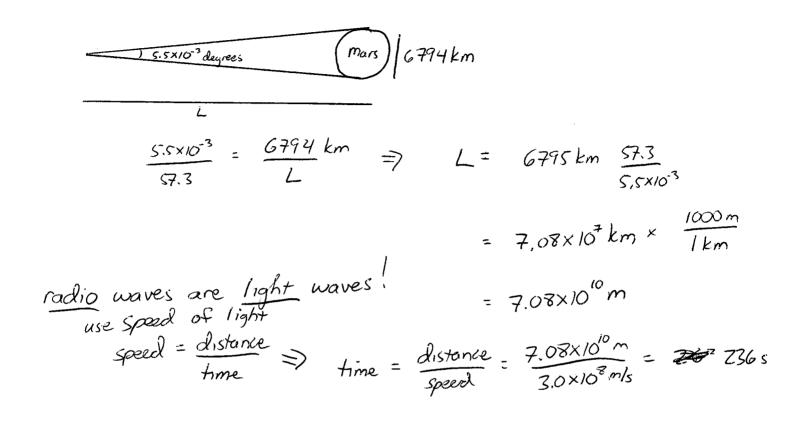
-member of Kuiper Belt (composition, orbit)

- not even largest kBO

- eccentric, inclined orbit

- mass less that all other planets and many moons ASTR 101 Third Hour Exam

SA2. Mars, whose diameter is 6794 km, has an angular size of 5.5×10^{-3} degrees when viewed from Earth this week. How long does it take radio signals from the Spirit and Opportunity rovers on the surface of Mars to reach the NASA scientists here on Earth?



SA3. Explain why scientists think that environmental conditions were different on Mars some 2 billion years ago. Make sure you explain how things were different as well as what evidence leads us to this conclusion.

evidence: dry channels on surface => running water in past extinct volcances => active vulcanism in past.

Running surface water implies that Mars once had a thicker, warmer atmosphere. Active vulcanism provided gaseous input, thickening the atmosphere. With a thicker atmosphere, the greenhouse effect warmed the surface. It with higher atmospheric pressure to warmer temps, wo water surface. It with higher atmospheric pressure to warmer temps, wo water can exist in higher form (e.g. won't evaporate away or freeze)

Problems

Show your work! I will be very generous with partial credit if I can figure out what you're doing!

Here is some information regarding Pluto and it's moon Charon:

Property	Pluto	Charon	
Radius	1160 km	$635~\mathrm{km}$	_
Mass	?	$1.8 \times 10^{21} \; \mathrm{kg}$	
Orbital radius	$5.9 imes 10^9 \; \mathrm{km}$	$1.96 imes 10^4 \; \mathrm{km}$	
(assume circular orbits)			
Orbital period	$7.83 \times 10^9 \mathrm{\ s}$	$5.51 \times 10^5 \mathrm{\ s}$	

P1. Calculate Charon's average density.

(20 points)

Density =
$$\frac{mass}{volume}$$

Volume =
$$\frac{4}{3}\pi radius^3$$
 $r = 635 km \times (1000 m) = 635000 m$
= $\frac{4}{3}(3.14)(637,000 m)^3$
= $1.07 \times 10^{18} m^3$

Density =
$$\frac{1.8 \times 10^{21} \text{kg}}{1.07 \times 10^{18} \text{ m}^3} = 1680 \text{ kg/m}^3$$

P2. Calculate Pluto's mass.

(20 points)

224 m/c

speed of Choron =
$$\frac{2\pi r_{orb}}{P}$$
 Charon's orbital properties
$$= \frac{2\pi (196 \times 10^7 \text{ m})}{5.31 \times 10^5 \text{ s}}$$

accel Charon =
$$\frac{v^2}{R} = \frac{(724 \, m/s)^2}{1.96 \times 10^3 \, m} = \frac{7.6 \times 10^{-3} \, m/s^2}{1.96 \times 10^3 \, m}$$

accelchavon =
$$\frac{GMpluto}{R^2}$$
 = $\frac{Mpluto}{R^2}$ = $\frac{accelchavon}{G}$ $\frac{R^2}{G}$

Charon & Pluto = $\frac{2.6\times10^{-3} \text{ m/s}^2}{G}$ (1.96)

$$= \frac{2.6 \times 10^{-3} \, \text{m/s}^2 \, \left(1.96 \times 10^3 \, \text{m}\right)^2}{6.67 \times 10^{-11} \, \text{m}^3/\text{kg s}^2}$$